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INSTITUTE



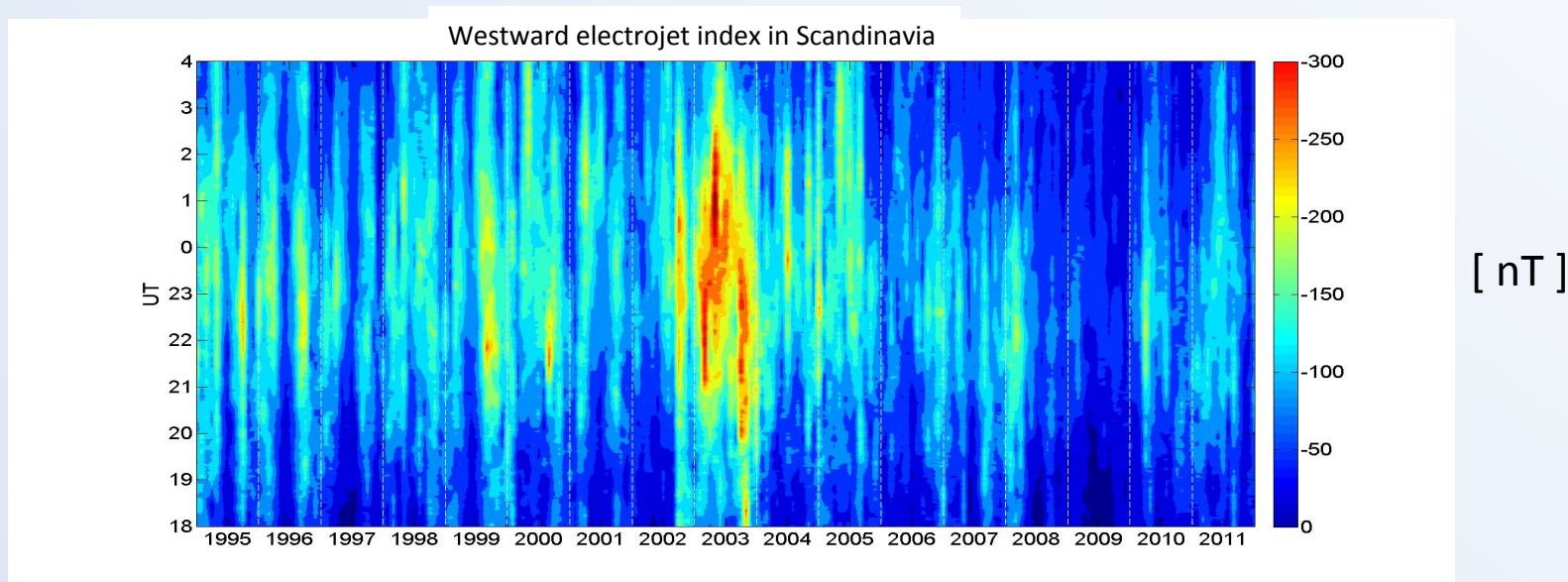
Aalto University

# Solar cycle evolution of Alfvénic fluctuations and their geo-efficiency

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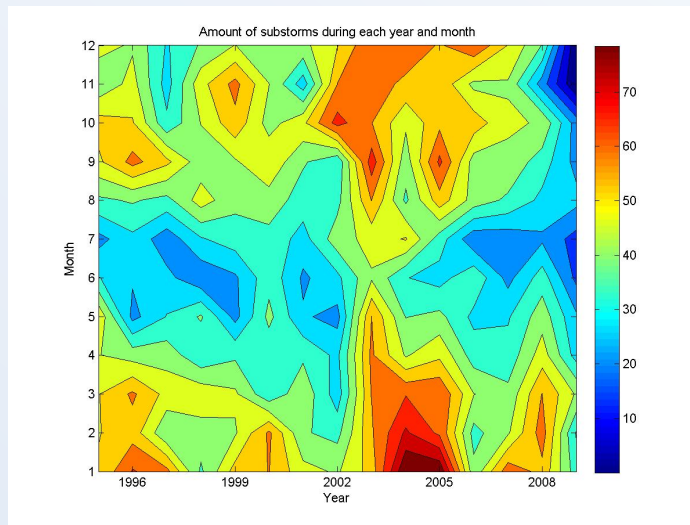
- (1) ReSoLVE Centre of Excellence, Finnish Meteorological Institute, (2) Aalto-university,  
(3) Birkeland Centre of Excellence, University of Bergen, Norway (4) Oulu university,  
(5) UCL/MSSL, UK, (6) Michigan university, Ann Arbor, US, (6) NASA Goddard SFC, US.

- Question 1:  
How does solar wind Alfvénicity evolve over the solar cycle 23?
- Question 2:  
Which solar wind structures carry most of the Alfvénicity?
- Questions 3: How does the solar wind Alfvénicity affect high-latitude geomagnetic activity in particular substorm activity.

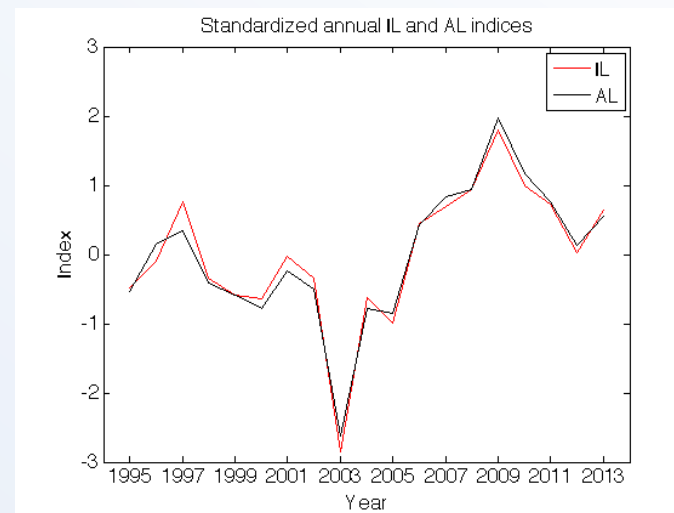


2003

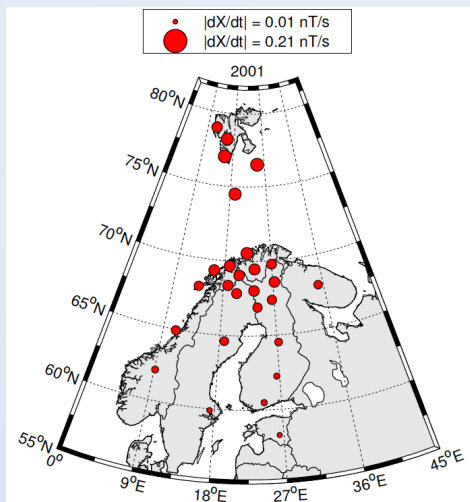
## Substorms



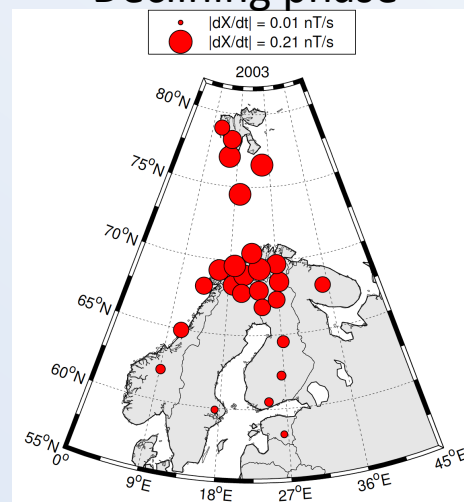
## IL and AL indices



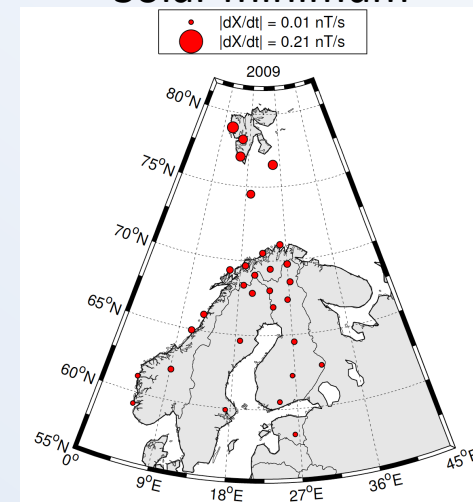
## Solar maximum



## Declining phase

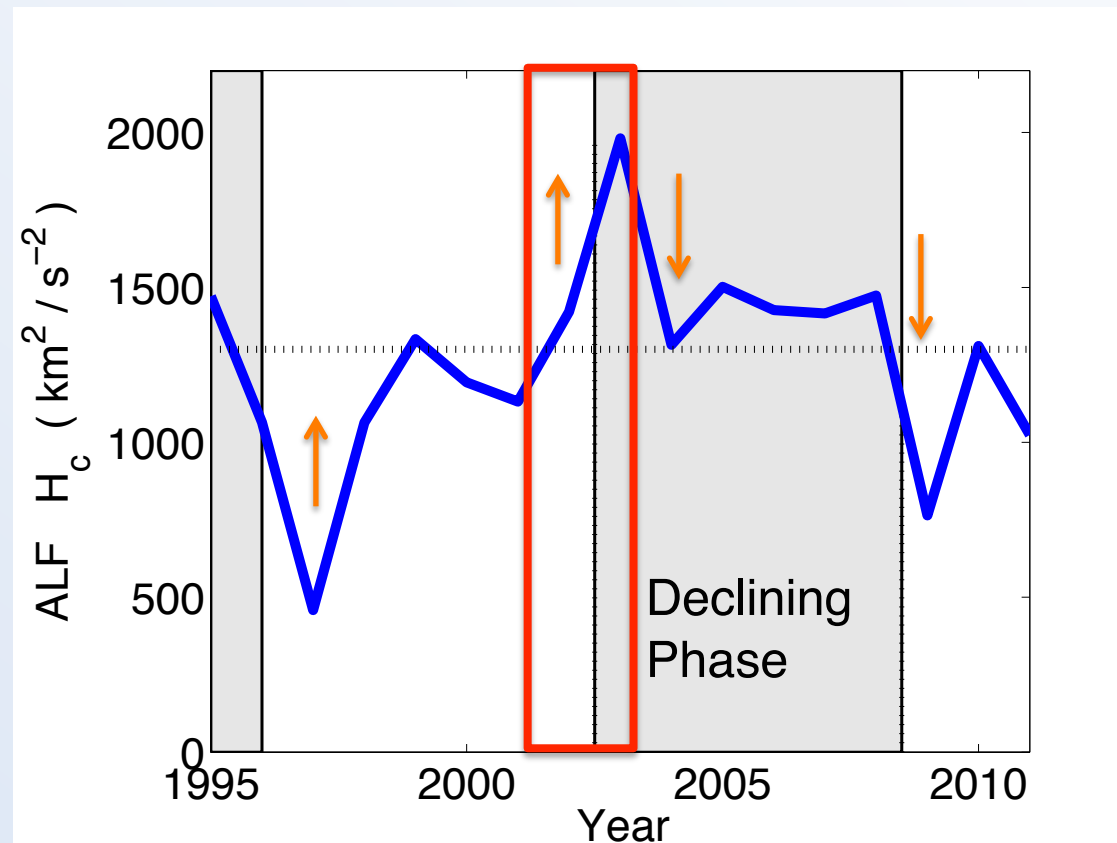


## Solar minimum



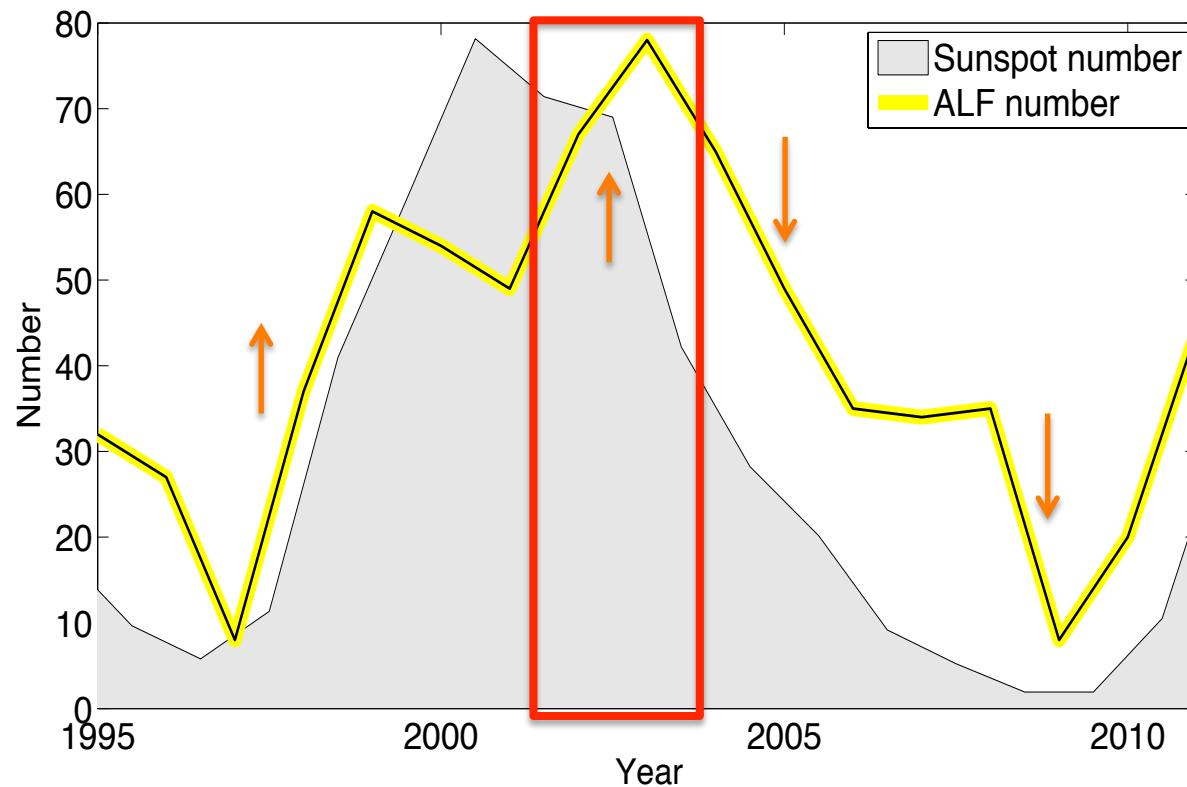
- We examine geomagnetic activity (IL and AL index) and substorms identified from them.
- Substorms identified from rapid decrease of north-south component of ground magnetic field and storms when Dst index  $< -40$  nT.
- High-speed streams  $v > 600$  km/s over 6 h.
- Alfvénic solar wind fluctuations (ALFs) identified based on normalized cross helicity by using method developed by Tu and March (1995) and further modified by Snekvik et al., 2013.

A rapid increase of Alfvénicity in early declining phase when solar wind Alfvénicity increase 40% from 2002 to 2003.



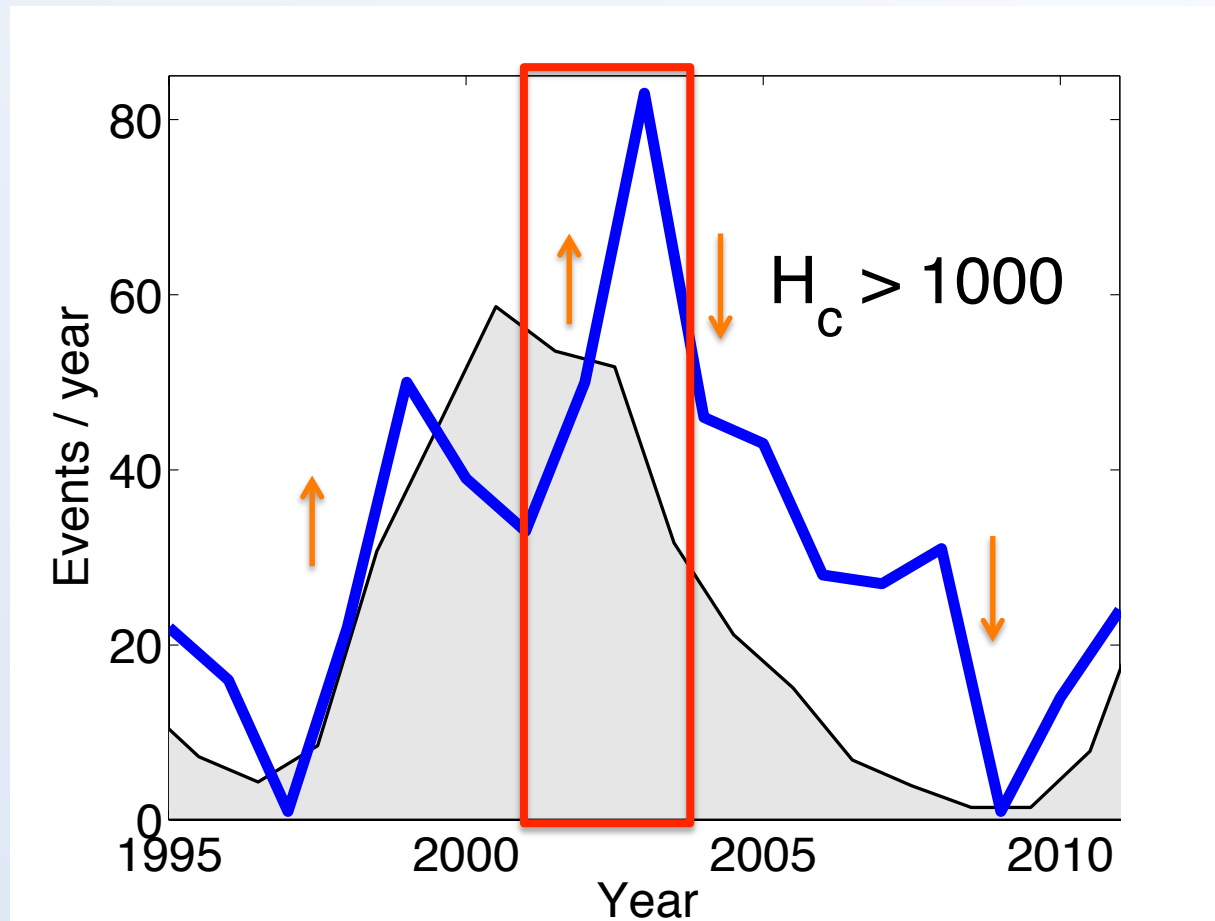
Tanskanen, Snekvik, Mursula et al., JGR, 2016.

Alfvénic fluctuations are found in all solar cycle phases, but largest ALF number is found in early declining phase in 2003.



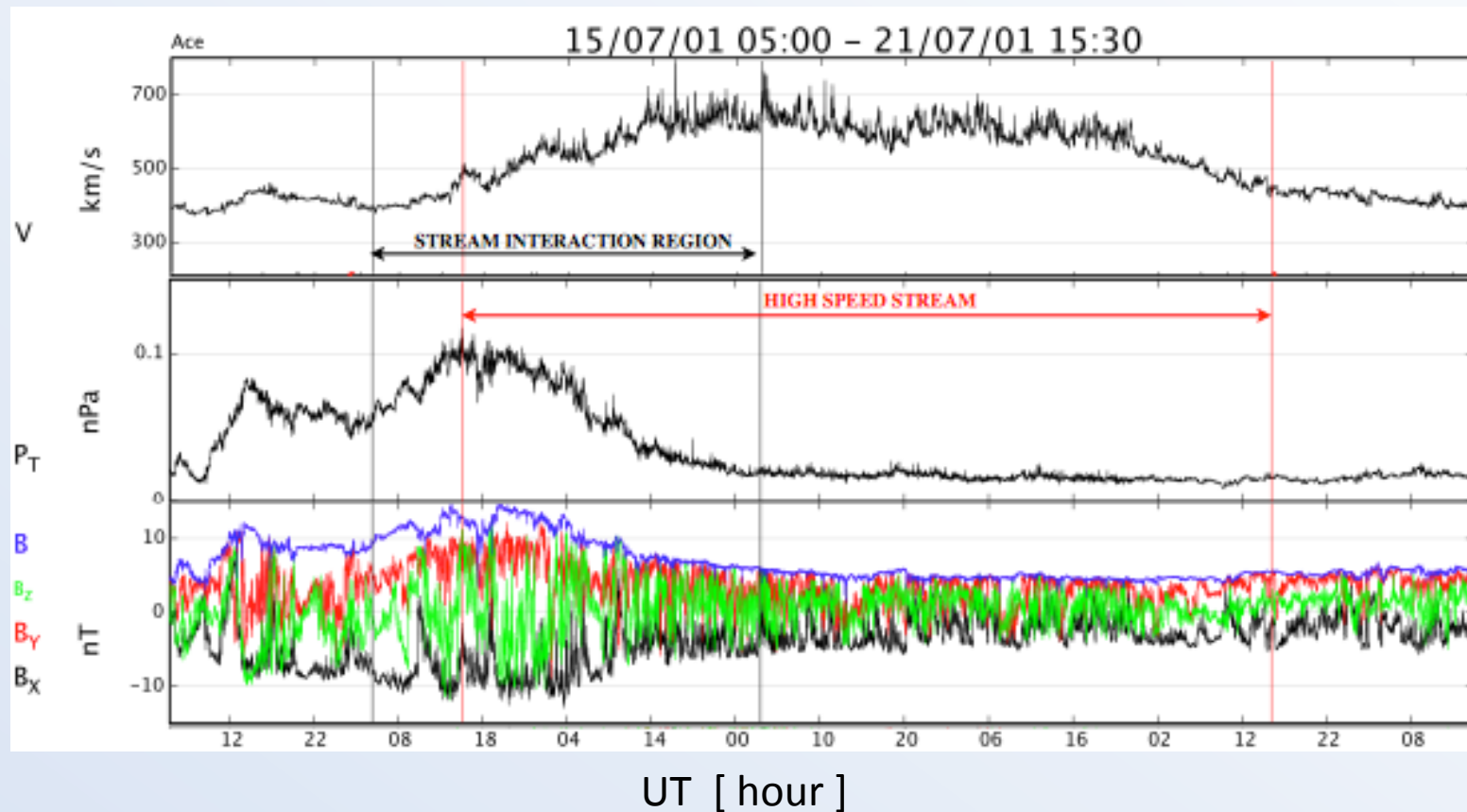


Extremely Alfvénic solar wind in 2003 and secondary peaks in 1999 and 2008.



Number of high-speed streams each year, when 6 h and 700 km/s criteria used:

	1993	94	95	96	97	98	99	00	01	02	03	04	05	06	07	08	09	10
HSS	2	21	4	0	0	3	2	8	5	7	27	3	13	1	0	2	0	1

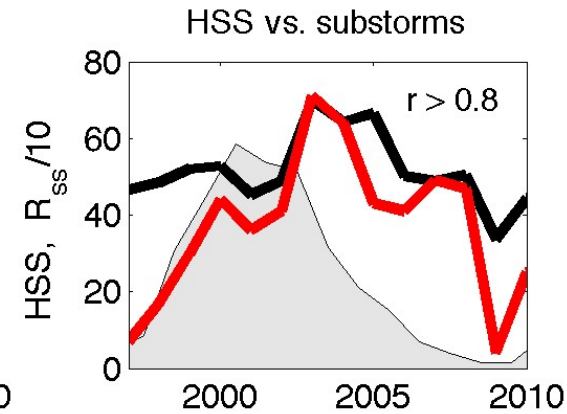
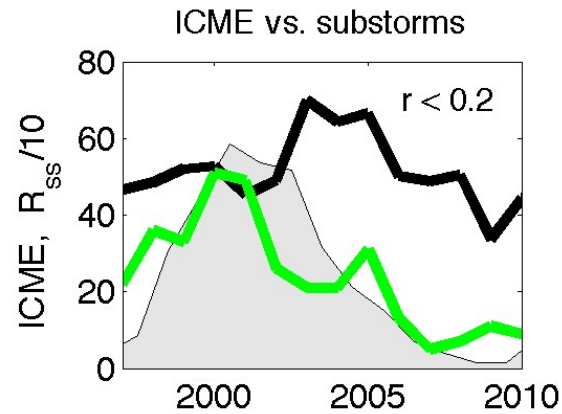




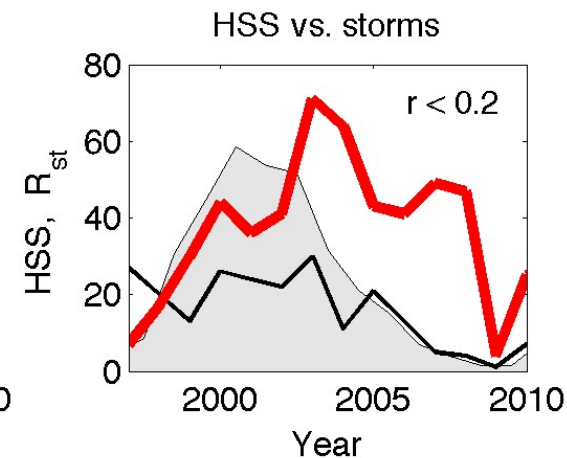
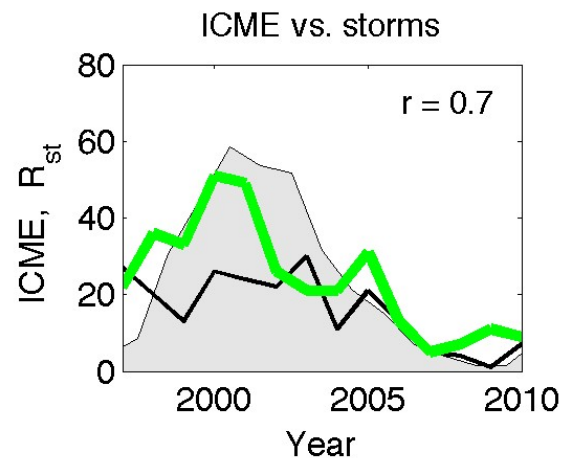
## Coronal mass ejections

## High-speed streams

Substorms



Storms

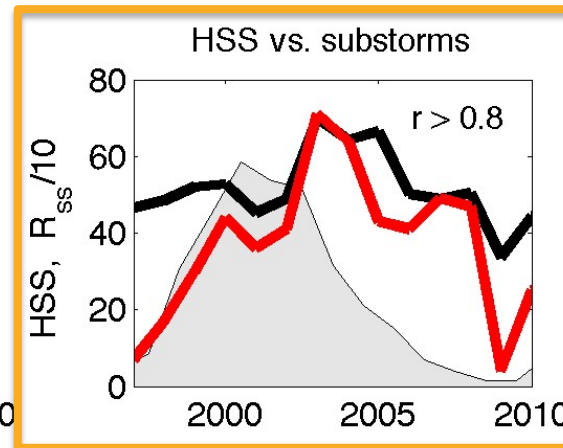
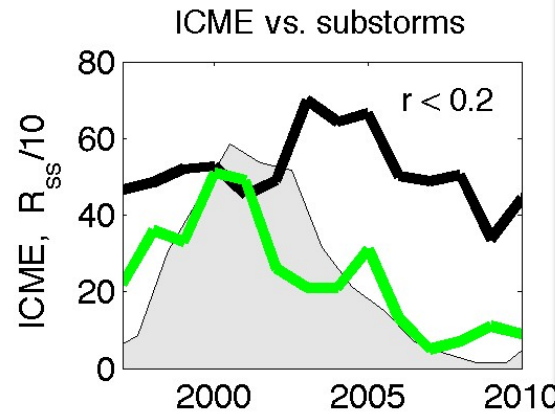


... and ICMEs modulate low latitudes.

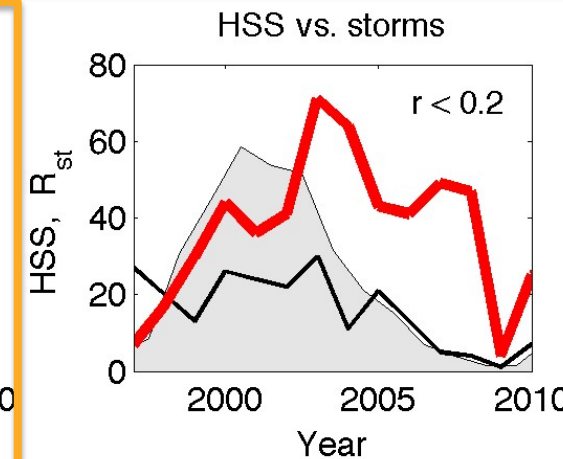
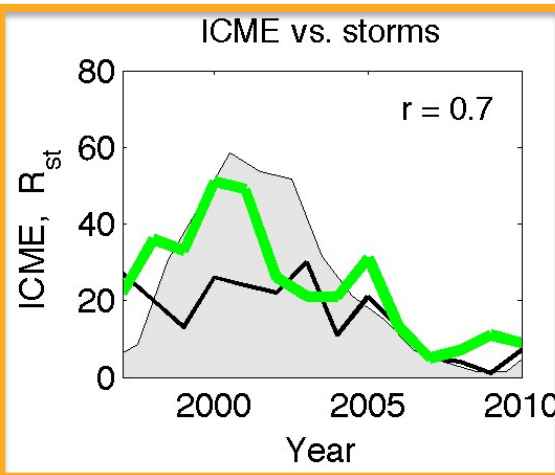
## Coronal mass ejections

## High-speed streams

Substorms

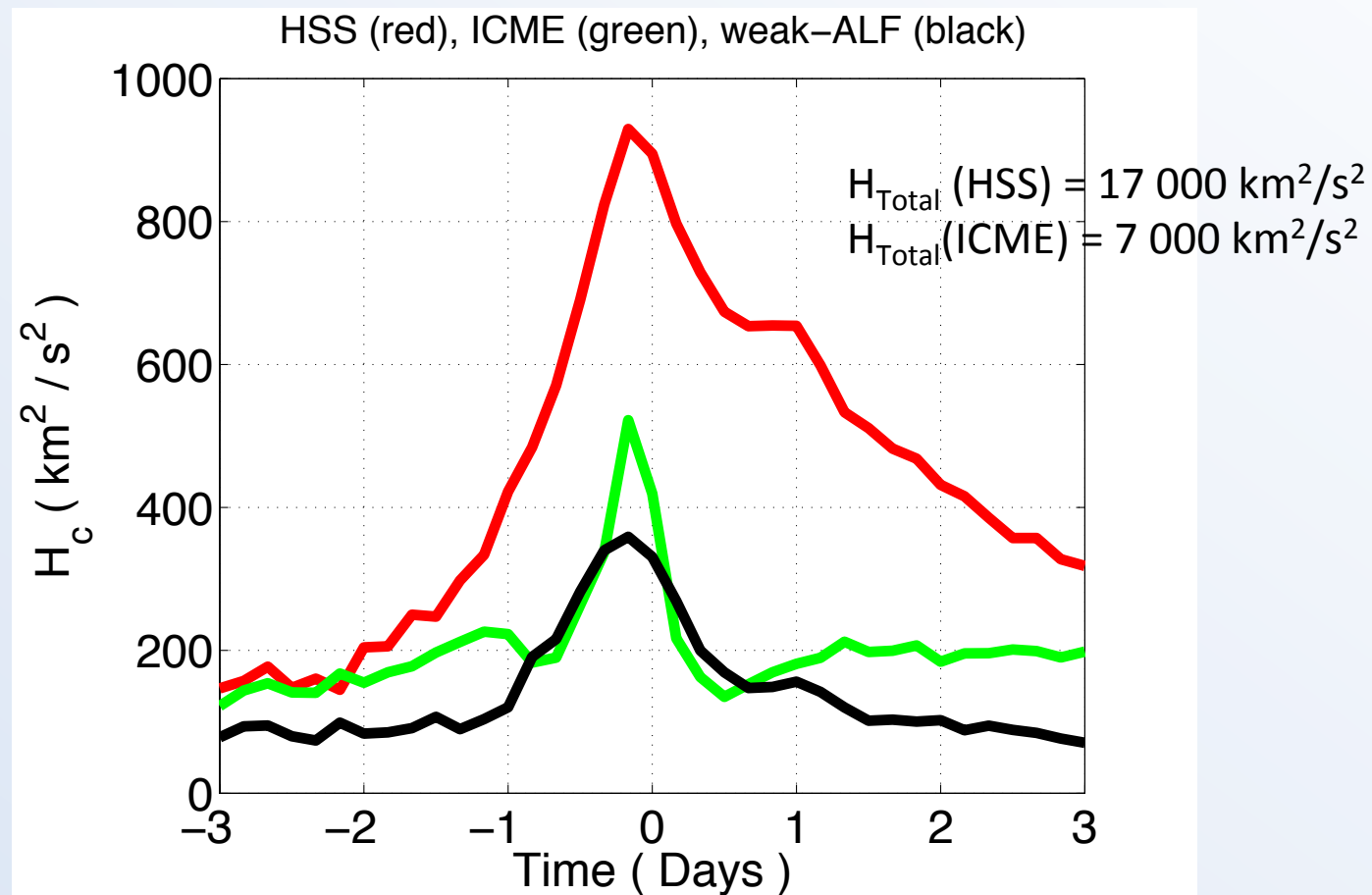


Storms

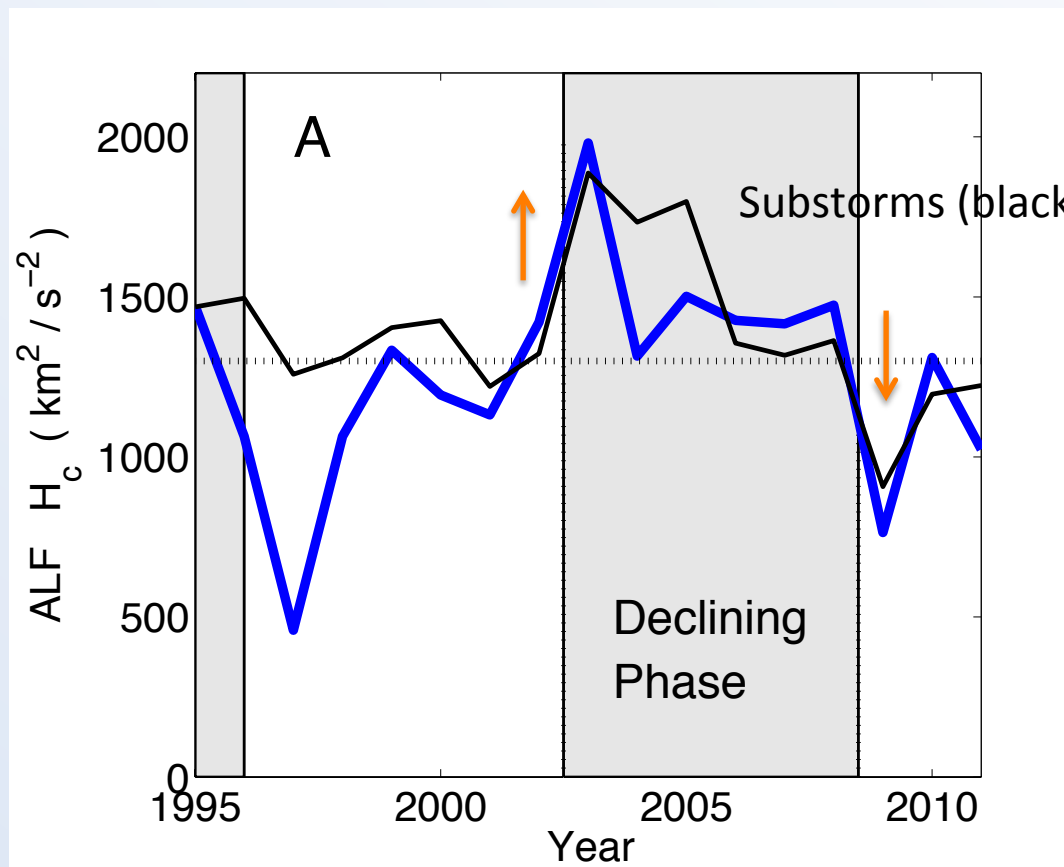


... and ICMEs modulate low latitudes.

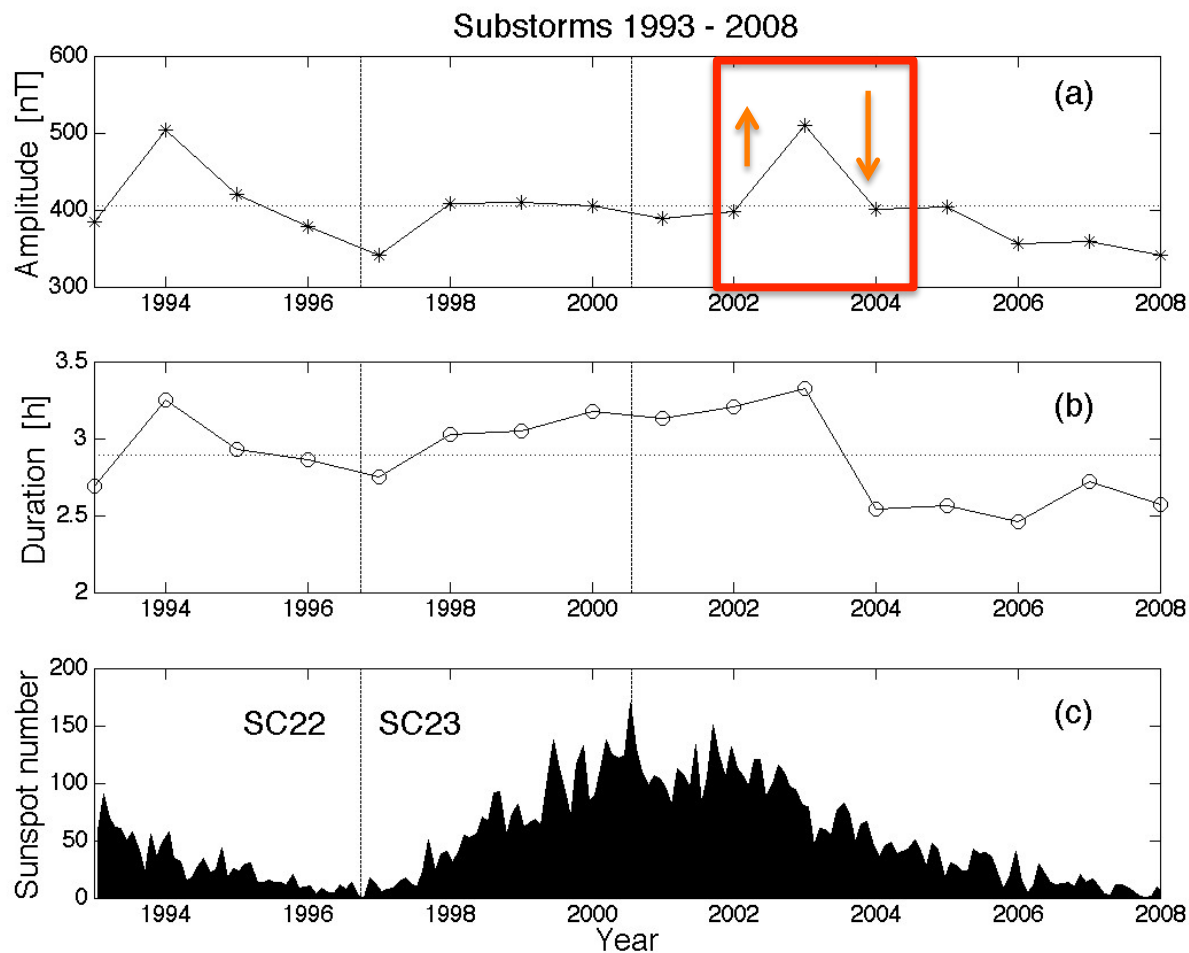
Alfvénic fluctuations carry 2.4 times more helicity during HSS than ICME -related solar wind.

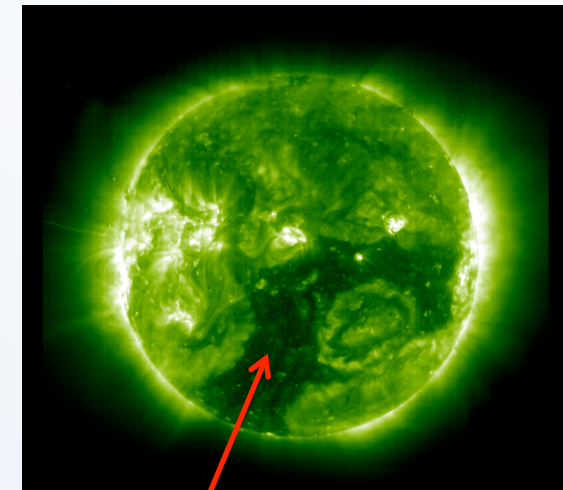
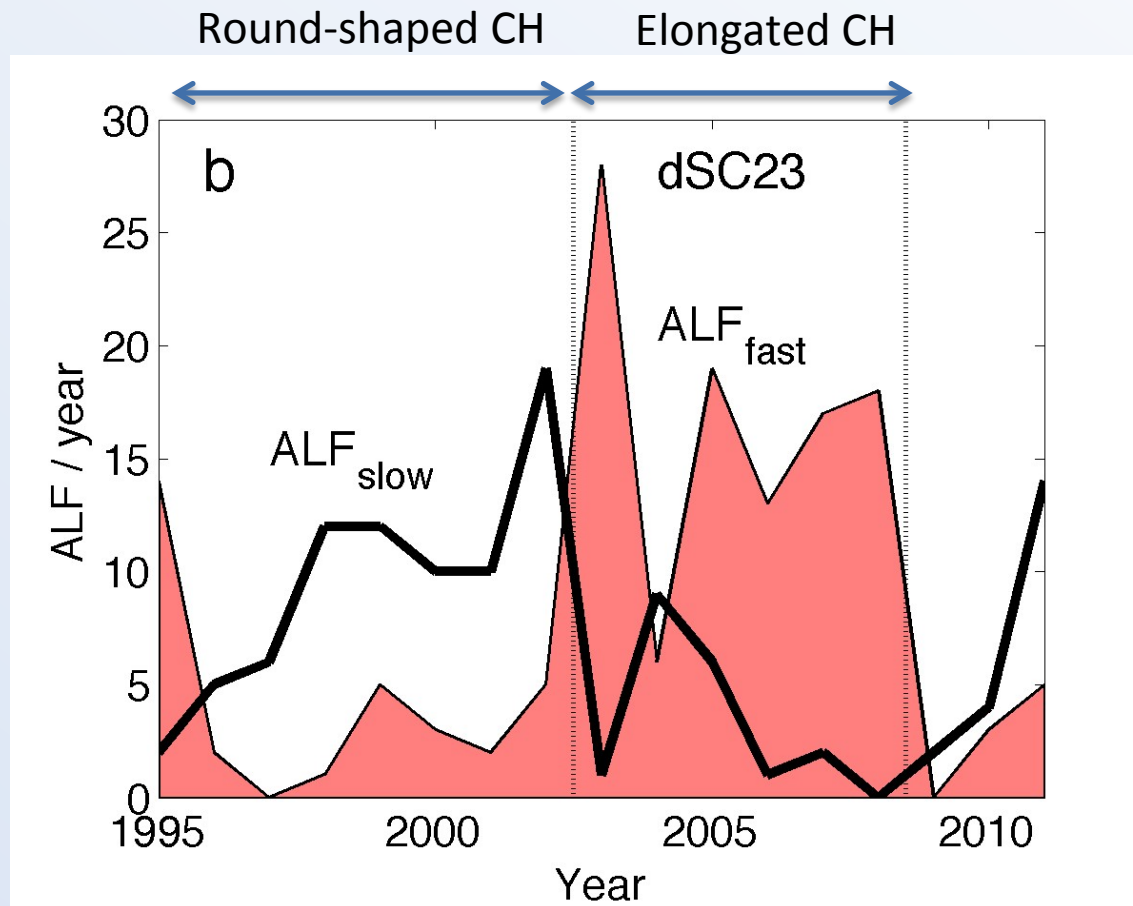


- \* An increase of 40 % in substorm number from 2002 to 2003 and 33 % decrease from 2008 to 2009.
- \* Alfvénicity increased 40 % from 2002 to 2003 and decreased 34 % from 2008 to 2009.

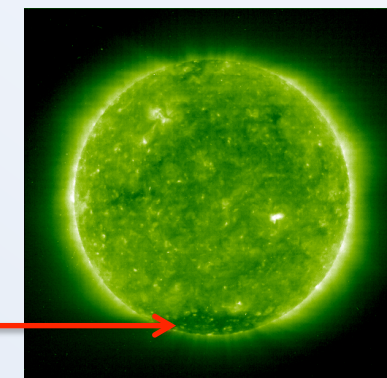


Substorm intensity increased 35% and decreased 25 % from 2003 to 2004, while Alfvénicity increased 40 % from 2002 to 2003 and decreased 34% from 2003 to 2004 .





Fast wind from elongated coronal holes.



Round-shaped coronal holes.



Characteristics of the solar wind and the interplanetary space changed in late 2002 when the solar maximum of the SC23 turned into the declining phase. After the transition:

- (1) High-speed streams (HSS) occur more frequently.
- (2) HSS embedded Alfvénic fluctuations are more intense and faster.
- (3) High-latitude geomagnetic activity is enhanced e.g. substorms occur more often and are more intense.

- Q1: How does solar wind Alfvénicity evolve over the solar cycle 23?  
Largest Alfvénicity is observed in early declining solar cycle phase in 2003 and in velocity range 550 – 600 km/s
- Q2: Which solar wind structures carry most of the Alfvénicity?
  - HSS carry twice as much Alfvénicity as ICMEs, and the largest Alfvénicity is seen in early trailing part of HSS.
- Q3: How does the solar wind Alfvénicity affect high-latitude geomagnetic activity in particular to substorm activity.
  - The 40% increase in the strength of ALFs from 2002 to 2003 increase substorm amplitude roughly by 30% and the substorm number roughly by 40%.